



ϕ, λ - lat & long of observer

δ - declination of the celestial object.

z - zenith angle of the object (would also need ^{local}azimuth to locate it in the sky).

The unit vector along the zenith vector is

$$\mathbf{r} = \begin{bmatrix} \cos\phi \cos\lambda \\ \cos\phi \sin\lambda \\ \sin\phi \end{bmatrix}$$

The unit vector along the celestial object is

$$\mathbf{u} = \begin{bmatrix} \cos\delta \cos t_0 \\ -\cos\delta \sin t_0 \\ \sin\delta \end{bmatrix}^T$$

The zenith angle is $\cos z = \mathbf{r}^T \mathbf{u}$.

Suppose the zenith angles to two objects are

$$r^T u_1 = \cos z_1, \quad r^T u_2 = \cos z_2.$$

This implies $r^T (\cos z_2 u_1 - \cos z_1 u_2) = 0$.

Let $h = \cos z_2 u_1 - \cos z_1 u_2$ & $g = u_1 \times u_2$. Note that $g \perp h$.

Since $r^T h = 0$, r lies in the plane formed by g & $g \times h$.

∴ Suppose $r = a_1 g + a_2 g \times h$.

Since $r^T r = 1$, we have $a_1^2 g^T g + a_2^2 g^T g \times h^T h = 1$.

Also, ~~$\cos z_1$~~ $= r^T u_1 = a_2 (g \times h)^T u_1$
 $= a_2 [(b u_1 \times b u_2) \times (\cos z_2 u_1 - \cos z_1 u_2)]^T u_1$
 $= -a_2 \cos z_1 (u_2 \times u_1)^T (u_1 \times u_2)$
 $= -a_2 \cos z_1 g^T g$.

∴ $a_2 = 1/g^T g$, which yields

$$a_1 = \pm \sqrt{\frac{g^T g - h^T h}{g^T g}}$$

∴ $r = \frac{1}{g^T g} [g \times h \pm \sqrt{g^T g - h^T h} g]$ — position fix.